## Towards the Theory of

 $Quantum\ Relativity$ 

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# 人之能知天也, 能知人之天也。

~王夫之

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Symmetry of "Space -Time-U...": -

the New Arena for Fundamental Physics
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Galilean Relativity

Einstein Relativity — special and general

(Special) Quantum Relativity

or Doubly (Triply) / Deformed Special Relativity
H.Synder / C.N.Yang (1947)

G.Amelino-Camelia (2001)

 $L.Smolin\,+\,(2002)$ 

····· General Quantum Relativity  $\Longrightarrow$  Quantum Gravity +

## My Basic Perspectives :-

- on issues related to Quantum Space-Time
- need new understanding of "Space-Time" (relativity principle)
- within Space-Time Geometry

  Non-Commutative Geometry is to Quantum Gravity

  as Non-Eucliean Geometry is to (Einstein) Gravity
- Quantum Field Theory, on (Minkowski/Einstein) Space-Time, perspective would be valid way beyond the experimentally probed energy/length scale a faith I don't share (Penrose)
- Linear (Vs Nonlinear) Realization
- an exploration into **new** ways to think about fundamental physics
- withhold phenomenological discussions

### A Stable Symmetry:-

R.V. Mendes (1994) Chryssomalakos & Okon (2004)

- perturbations (deformations) of a symmetry algebra
- stable if deformed algebras isomorphic to original
- structure constants perturbation Vs experimental uncertainty
- Examples
  - 1/ Galilean to Einstein Relativity  $(1/c^2 \neq 0)$
  - 2/ classical to quantum mechanics ( $\hbar \neq 0$ )

Poisson bracket  $\implies$  Moyal bracket

 $\bullet$  deformation parameters  $\longrightarrow$  fundamental constants

## Space-Time + Quantum Symmetry

'Poincaré + Heisenberg' algebra

$$\Longrightarrow$$
 SO(1,5) OR SO(2,4)

### WI Contraction:-

$$SO(1,3) \longrightarrow ISO(3) \ [N_i,N_j] = -i\,rac{1}{c^2}\,M_{ij}$$

 $\frac{1}{c^2} \to 0$  boosts commute (i.e. Galilean)

ullet SO(1,3) as the stablization :

$$[M_{ij}, L_{0k}] = i \left( \eta_{jk} L_{0i} - \eta_{ik} L_{0j} \right)$$

$$[L_{0i}, L_{0j}] = -i M_{ij}$$
take  $N_i \to \frac{1}{c} L_{0i}$ : isomorphic

? value of  $c: 3 \times 10^8 \text{ Vs} 1 \text{ or } 3 \times 10^{-7} \text{ (km ps}^{-1)}$ 

## Quantum Relativity SO(1,5) or SO(2,4):-

$$[J_{AB}, J_{LN}] = i (\eta_{BL} J_{AN} - \eta_{AL} J_{BN} + \eta_{AN} J_{BL} - \eta_{BN} J_{AL})$$

with  $\eta^{AB} = (1, -1, -1, -1, -1, \mp 1)$ .

- linear realization on 6-geometry  $J_{AB} = i (x_A \partial_B x_B \partial_A)$
- $J_{AB}$  for  $0 \to 3$  are the 10 generator for Lorentz symmetry  $M_{\mu\nu}$
- NOT necessarily means extra space-time dimensions
- 3 deformations  $ISO(3) \rightarrow SO(1,3) \hookrightarrow ISO(1,3)$  $\rightarrow SO(1,4) \hookrightarrow ISO(1,4) \rightarrow SO(1,5) \text{ or } SO(2,4)$
- 3 invariant quantities : c  $\kappa$   $\ell$  Kowalski-Glikman & Smolin(2004) OK (2007)
- ◆ curved momentum space ⇒ non-commutative space-time
   + curved space-time ⇒ non-commutative momentum space

#### The Three Deformations:-

$$SO(1,4) :- |z^A| \le i \ell$$
  
 $\eta_{AB} z^A z^B = -\ell^2 \left(1 + \frac{1}{G^2}\right)$ 

$\Delta x^i(t) = \frac{\mathbf{v}^i}{\mathbf{v}^i} \cdot t$	$\Delta x^{\mu}(\sigma) = \mathbf{p}^{\mu} \cdot \sigma$	$\Delta x^A(\rho) = \mathbf{z}^A \cdot \rho$
$ v^i  \le c$	$ p^{\mu}  \le \kappa c$	$ z^A  \le \ell$
$-\eta_{ij}v^iv^j = c^2\left(1 - \frac{1}{\gamma^2}\right)$	$\eta_{\mu\nu}p^{\mu}p^{\nu} = \kappa^2 c^2 \left(1 - \frac{1}{\Gamma^2}\right)$	$\eta_{AB} z^A z^B = \ell^2 \left( 1 - \frac{1}{G^2} \right)$
$M_{0i} \equiv N_i \sim P_i$	$J_{\mu 4} \equiv O_{\mu} \sim P_{\mu}$	$J_{A5} \equiv O_A' \sim P_A$
$[N_i,N_j] \longrightarrow -iM_{ij}$	$[\mathit{O}_{\!\mu}, \mathit{O}_{\! u}] \longrightarrow i M_{\mu u}$	$[O_A', O_B'] \longrightarrow i J_{AB}$
$\vec{u}^4 = \frac{\gamma}{c}(c, v^i)$	$\vec{\pi}^5 = \frac{\Gamma}{\kappa  c}(p^\mu, \kappa  c)$	$ec{X}^6 = rac{G}{\ell}(z^A,\ell)$
$\eta_{\mu\nu}u^{\mu}u^{\nu}=1$	$\eta_{AB}\pi^A\pi^B=-1$	$\eta_{\mathcal{M}\mathcal{N}}X^{\mathcal{M}}X^{\mathcal{N}} = -1$
$IR^3 \rightarrow SO(1,3)/SO(3)$	$M^4 \rightarrow SO(1,4)/SO(1,3)$	$M^5 \rightarrow SO(1,5)/SO(1,4)$

- 4-momentum is defined by  $p^{\mu} = \frac{dx^{\mu}}{d\sigma}$
- $z^A = \frac{dx^{\mu}}{d\rho}$  is chosen as a length
- without the  $p^{\mu}$  deformation,  $X^{\mathcal{M}}$  are just alternative coordinates of a de-Sitter space-time
- ullet no  $\hookrightarrow ISO(1,5)/ISO(2,4)$  : no naive translation symmetries

## Momentum Boosts ( with $\hat{X}_{\mu}$ ) :-

$$J_{\mu 4} = i \left( x_{\mu} \partial_4 - x_4 \, \partial_{\mu} \right)$$

— rotations among  $x_{\mu}$ 's and  $x_4 = -\kappa c \sigma$ 

A.Das & O.K.

 $\star \sigma$  is peculiar — (doubt space-time interpretation!)

 $\bullet \ ISO(1,3) \qquad \to \ \ SO(1,4)$ 

since Snyder

- $\kappa$  as Planck mass
- preferred frame for a quantum state :  $\vec{\pi}^{\,5} = (0, 0, 0, 0, 1)$ the reference frame does not see its own 4-momentum/mass
- boost characterized by  $p^{\mu}$ , hence mass dependent
- quantum frames

- $p^{\mu} = \frac{dx^{\mu}}{d\sigma}$  Vs  $m c u^{\mu} = \gamma m c \frac{dx^{\mu}}{dx^{0}}$  (as Einstein/classical limit )  $\sigma \longrightarrow \frac{\tau}{m}$  for an Einstein particle
  recall: "space-like" geometrical signature
- general transformation on-shell condition not preserved classical to quantum frame, uncertainty: observer to observed
- $\sigma$ -coordinate Vs literature
- Vs proper time / covariant Schrödinger QM
- system with indefinite mass
- QM on  $M^4$  stochastic metric fluctuation
- \* momentum boosts needed for group quantization (relativistic)

#### Quantum Frame of Reference:-

Aharonov & Kaufherr, PR 30, 368 C. Rovelli, Class.Quantum.Grav. 8, 317

- context: nonrelativistic QM / gravitation
- characterized by mass (together with  $v + \cdots$ )
- interaction of frame (observer/device) with observed
- observation has nontrivial effect on observer
- uncertainty principle on frame (observer/device)
  - → uncertainty on observed
- grav. local gauge inv. observables obtained after dynamic properties of frame (observer/device) taken into consideration
- \* Is there such a thing as a practical quantum measurement?

## Translational Boosts ( with $P_a^{(\mathcal{L})}$ — $\hat{P}_{\mu}$ ) :-

$$J_{A5} = i \left( x_A \partial_5 - x_5 \, \partial_A \right)$$

— rotations among  $x_A$ 's and  $x_5 = -\ell \rho$ 

A.Das & O.K. O.K.

- $\star \rho$  may characterize the scale
- boost characterized by  $z^A$ , a location vector sort of a translation
- preferred frame for a state :  $\vec{X}^6 = (0, 0, 0, 0, 0, 1)$ the reference frame does not see its own location — (0, 0, 0, 0, 0, 1) characterizes the coordinate origin  $z^A = 0$
- \* a reference frame does not observe itself (no  $v^i$ ,  $p^{\mu}$ ,  $z^A = 0$ )
- Beltrami description *preserves* physics of 5-momentum constraint

A.Das & O.K.

• ISO(1,4)  $\to$  SO(1,5)  $-\ell$  an IF cutoff (Vs  $\kappa$ )  $\leftrightarrow$  cosmological constant

Kowalski-Glikman & Smolin(2004)

- $-dS_5$  as "quantum" world

O.K.

- $ullet ISO(1,4) 
  ightarrow SO(2,4) \ -\ell_P ext{ as Planck length (UV cutoff)}; \qquad \kappa \, c \, \ell_P = \hbar$  $(\kappa^2 = c \, \hbar/G \, , \, \ell_P^2 = \hbar \, G/c^3)$
- $AdS_5$  as quantum world
- ullet no ISO(1,5) or ISO(2,4) no coordinate translations
- quantum general relativity the role of G?

## Non-commutative (Space-Time) Geometry:-

cf. Kowalski-Glikman & Smolin Chryssomalakos & Okon

• (four) space-time position operators  $\hat{X}_{\mu} = -\frac{1}{\kappa c} i (x_{\mu} \partial_4 - x_4 \partial_{\mu})$ 

$$\hat{X}_{\mu} = -\frac{1}{\kappa c} i \left( x_{\mu} \partial_4 - x_4 \partial_{\mu} \right)$$

$$[\hat{X}_{\mu}, \hat{X}_{\nu}] = \frac{i}{\kappa^2 c^2} M_{\mu\nu}$$

(four) energy-momentum operators  $\hat{P}_{\mu} = \frac{1}{\ell} i (x_{\mu} \partial_{5} - x_{5} \partial_{\mu})$ 

$$\hat{P}_{\mu} = \frac{1}{\ell} i \left( x_{\mu} \partial_5 - x_5 \partial_{\mu} \right)$$

$$[\hat{P}_{\mu}, \hat{P}_{\nu}] = \frac{i}{\ell^2} M_{\mu\nu}$$

$$[\hat{X}_{\mu}, \hat{P}_{\nu}] = -i \, \eta_{\mu\nu} \hat{F} \,, \qquad [\hat{X}_{\mu}, \hat{F}] = +\frac{i}{\kappa^2 c^2} \hat{P}_{\mu} \,, \qquad [\hat{P}_{\mu}, \hat{F}] = -\frac{i}{\ell^2} \hat{X}_{\mu}$$

• 
$$i \partial_{\mu} \ll \kappa c$$
 and  $i \partial_{4} = p_{4} = -\kappa c$ :

$$\hat{X}_{\mu} \longrightarrow x_{\mu}$$

• 
$$x_{\mu} \ll \ell$$
 and  $x_5 = -\ell \ (\rho = 1)$ :  $\hat{P}_{\mu} \longrightarrow i \partial_{\mu} = p_{\mu}$ 

$$\hat{P}_{\mu} \longrightarrow i \, \partial_{\mu} = p_{\mu}$$

## SO(2,4)Quantum Relativity:-

—  $AdS_5$  as quantum world

• matching to 4D conformal symmetry:-

$$K_{\mu} \Rightarrow \sqrt{2}J_{\mu-}$$
  $P_{\mu} \Rightarrow \sqrt{2}J_{\mu+}$   $D \Rightarrow -J_{45}$  
$$J_{\mu\pm} \equiv i\hbar \left(x_{\mu}\partial_{\pm} - x_{\pm}\partial_{\mu}\right) = \frac{1}{\sqrt{2}}(J_{\mu 5} \pm J_{\mu 4})$$

• coordinate transformation — (gives the right matching)

$$(x^{\mu}, x^{4}, x^{5}) \longrightarrow (y^{\mu}, \frac{1}{2}\eta_{\mu\nu}y^{\mu}y^{\nu} + \frac{1}{2}, \frac{1}{2}\eta_{\mu\nu}y^{\mu}y^{\nu} - \frac{1}{2})$$

• However:  $\Rightarrow x_{+} = x^{-} = \frac{1}{\sqrt{2}}(x^{5} - x^{4}) = -\frac{1}{\sqrt{2}}$ 

$$x_- = x^+ = \frac{1}{\sqrt{2}} y^2$$
  $\partial_+ = 0$   $\partial_5 = -\partial_4 = \frac{1}{\sqrt{2}} \partial_- = x^{\nu} \partial_{\nu}$ 

\*  $\eta_{\mathcal{M}N} x^{\mathcal{M}} x^{\mathcal{N}} = \eta_{\mu\nu} x^{\mu} x^n u + 2 x_+ x_- = 0 \quad (\neq -1)$  for conformal universe \*  $x^5$ ,  $x^4$  (or  $x_- = x^+$ ) translations related to scaling?

#### Remarks:-

- (bottom line) an interesting, radical but sensible, approach
- primitive stage difficult to make (and identify) progress
- need creative but careful thinking about 'physics' beyond usual framework
- job for Einstein hope we can make minor steps "The chief cause of my failure was my clinging to the idea that the variable t only can be considered as the true time and my local t' must  $\cdots$  In Einstein's theory,  $\cdots$  t' plays the same part as t  $\cdots$ "

The Theory of Electrons (1916 ed.) — Lorentz (from A.I. Miller)

 ${\bf Space\text{-}Time\text{-}U}\;...$  as the New Arena for Fundamental Physics

## THANK YOU!